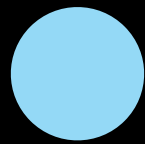
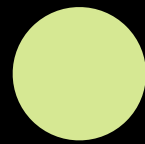
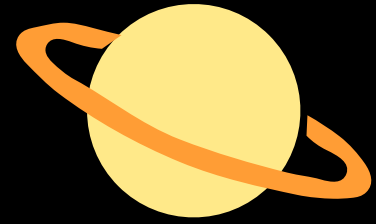


Hunting for habitable worlds

Elizabeth Tasker

Assoc. Professor
Institute of Space & Astronautical Sciences,
Japan Aerospace Exploration Agency (JAXA)





Sun

Mercury

Venus

Earth

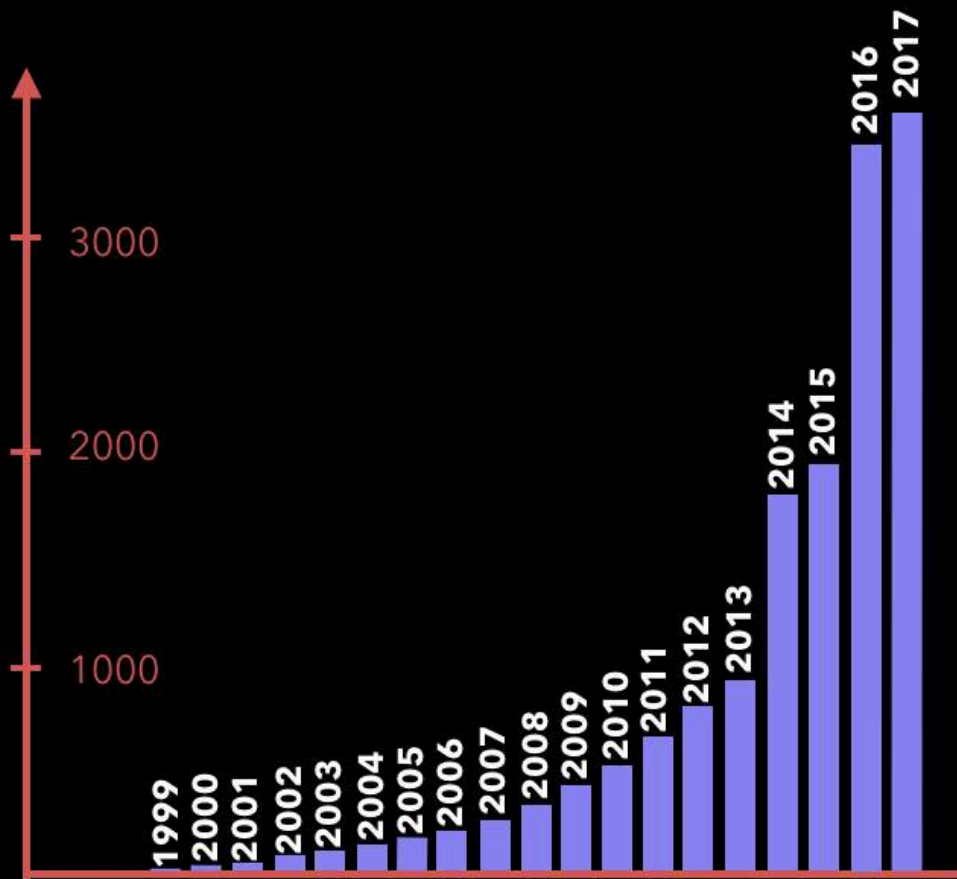
Mars

Jupiter

Saturn

Uranus

Neptune

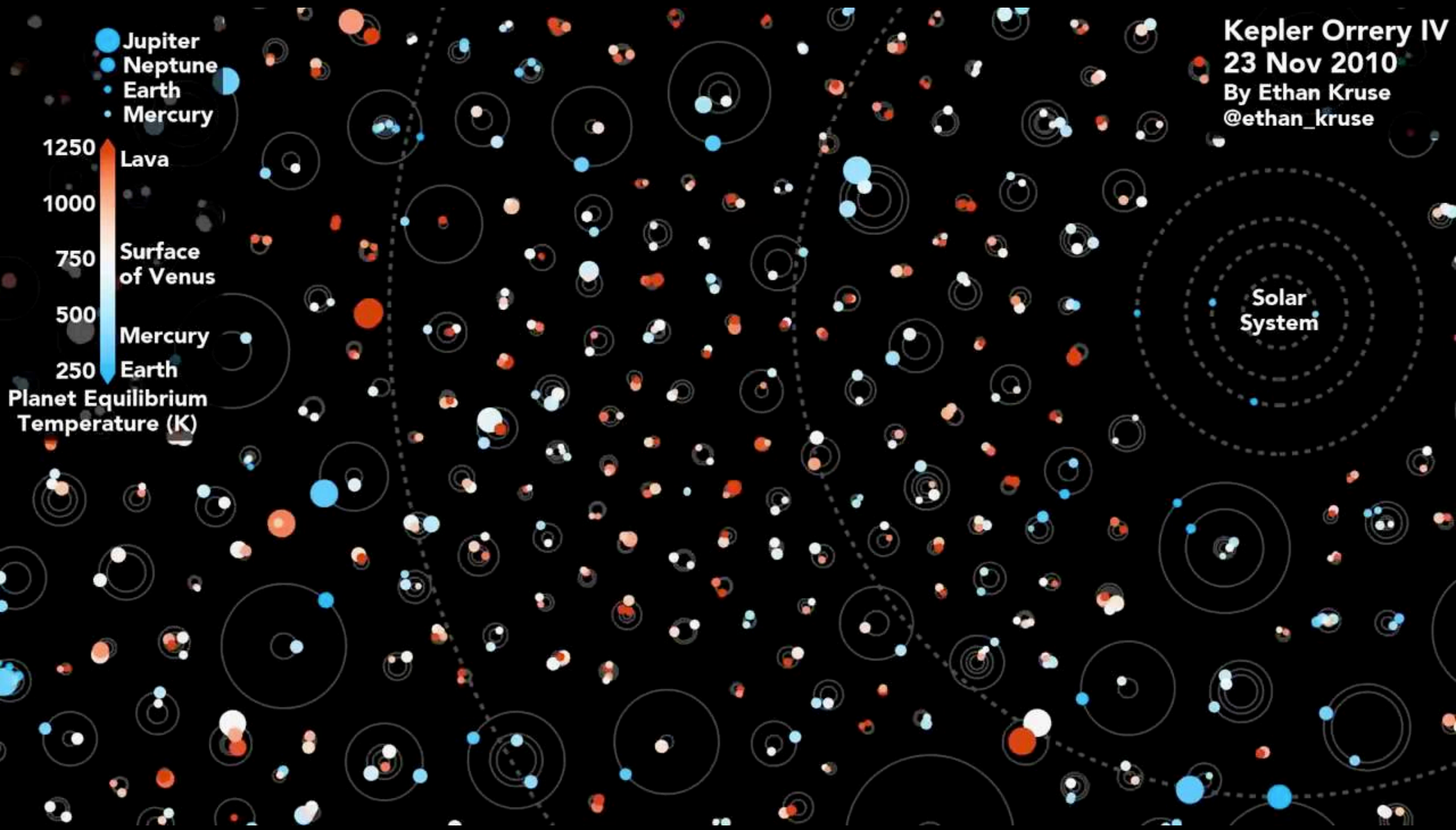


We now know of thousands of exoplanets:

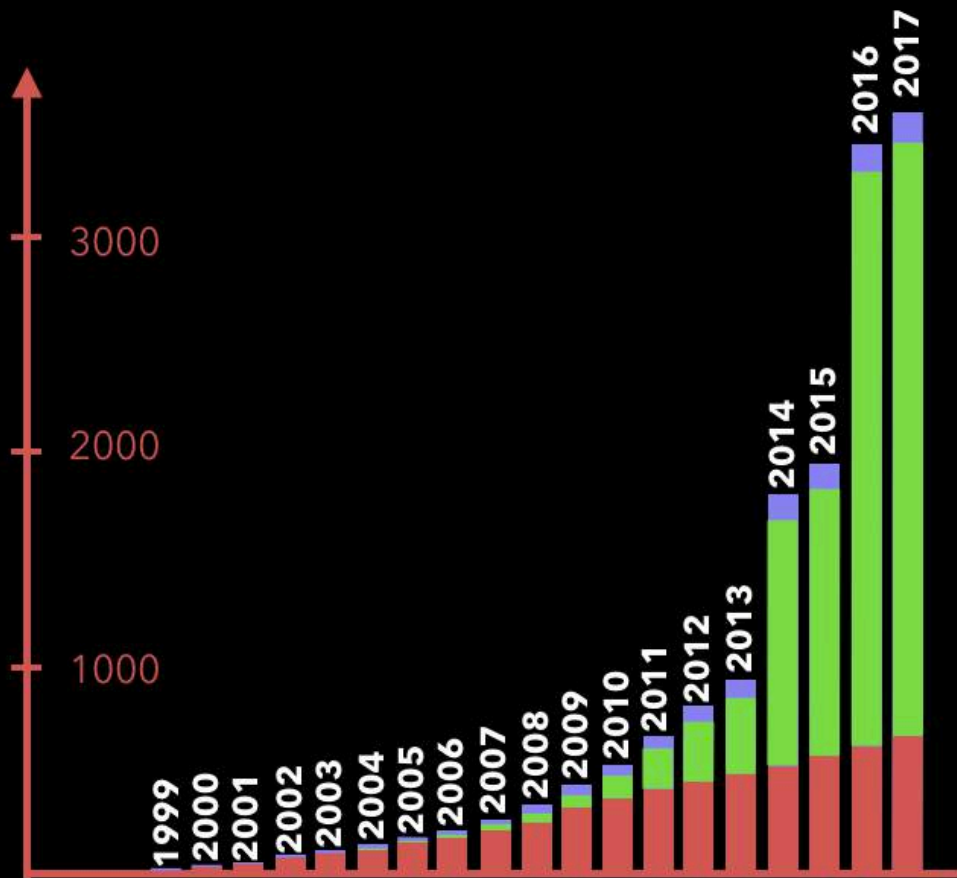
Planets that orbit stars beyond the Sun.

<https://exoplanetarchive.ipac.caltech.edu>

Kepler Orrery IV
23 Nov 2010
By Ethan Kruse
@ethan_kruse

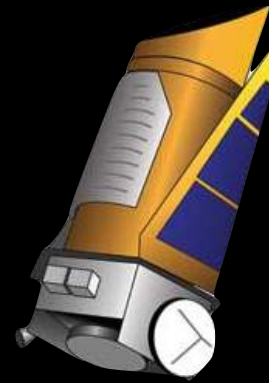


Could any of these new worlds be habitable?



<https://exoplanetarchive.ipac.caltech.edu>

- Transit
- Radial velocity
- Other

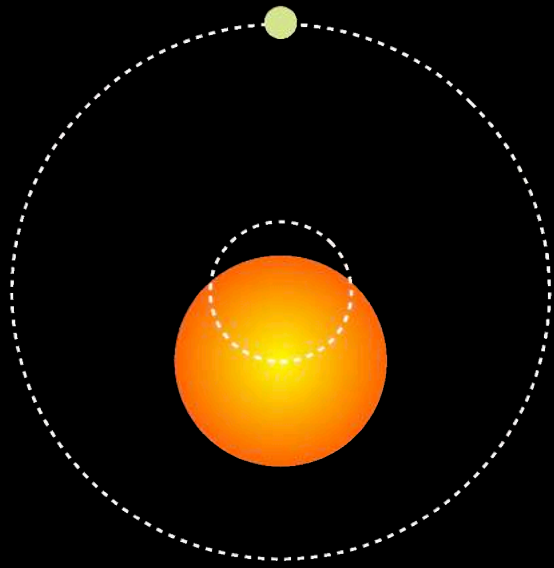


Kepler
(Transit)



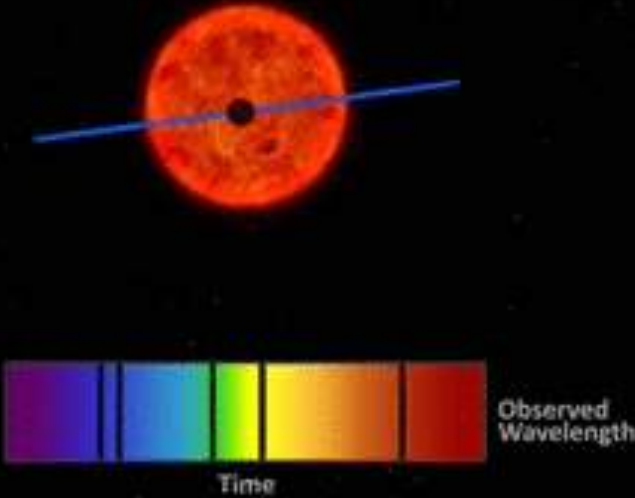
ESO La Silla Observatory
(Radial velocity)

96% exoplanets found by either the radial velocity or transit technique



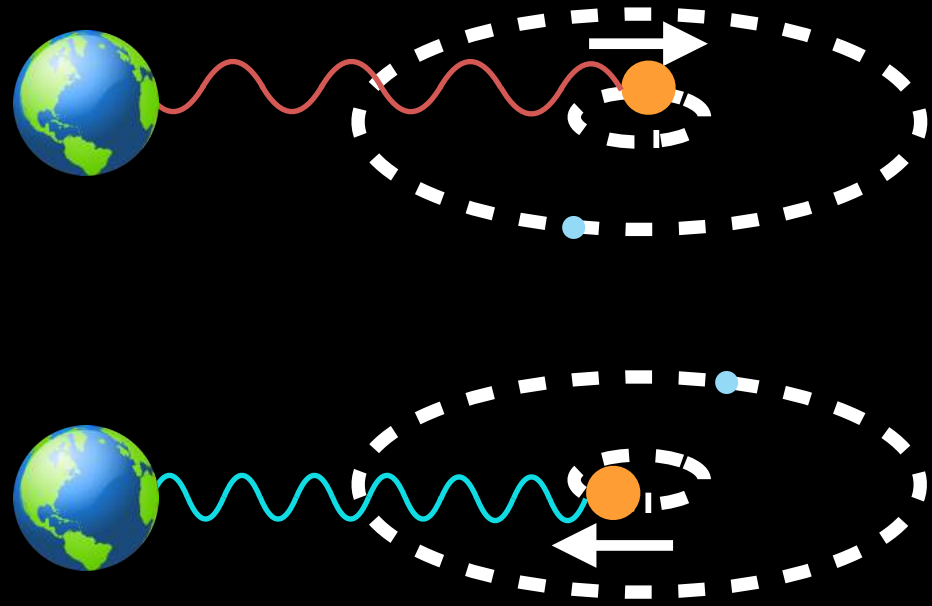
Radial velocity or **Doppler wobble**

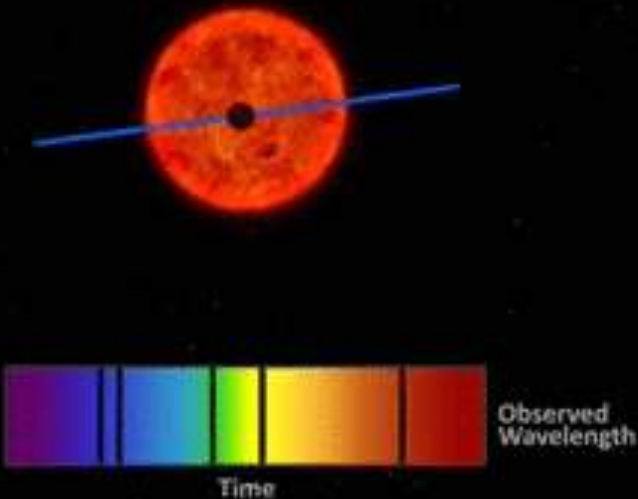
Orbit with planet causes the star to wobble, creating a periodic shift in wavelength



Radial velocity or Doppler wobble

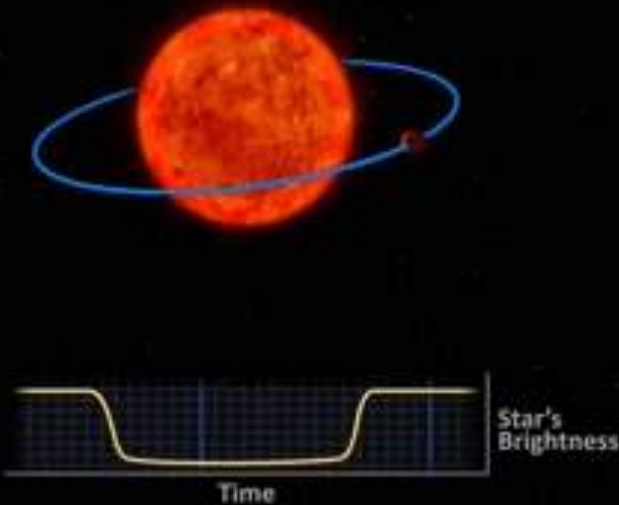
Orbit with planet causes the star to wobble, creating a periodic shift in wavelength





Radial velocity or Doppler wobble

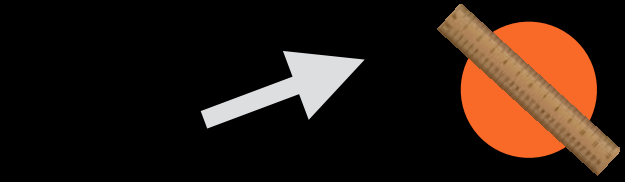
Orbit with planet causes the star to wobble, creating a periodic shift in wavelength



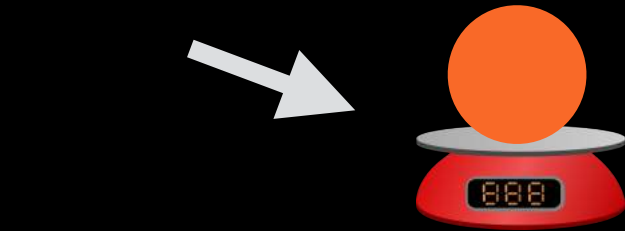
Transit

Dip in light as planet crosses our line of sight to the star

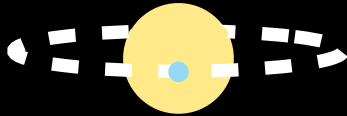
Typically, < 2 planet properties:



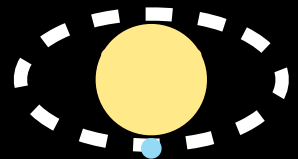
Planet radius



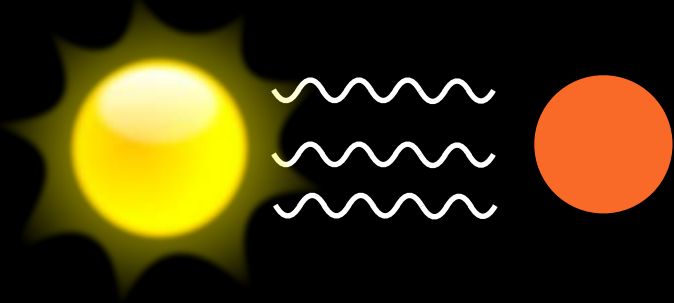
Planet (minimum) mass



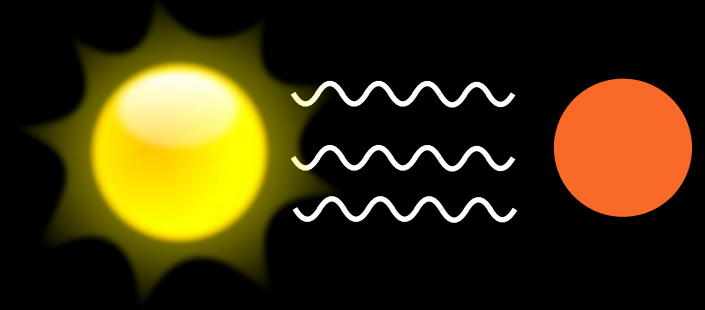
$m \sin(i) = \text{mass}$



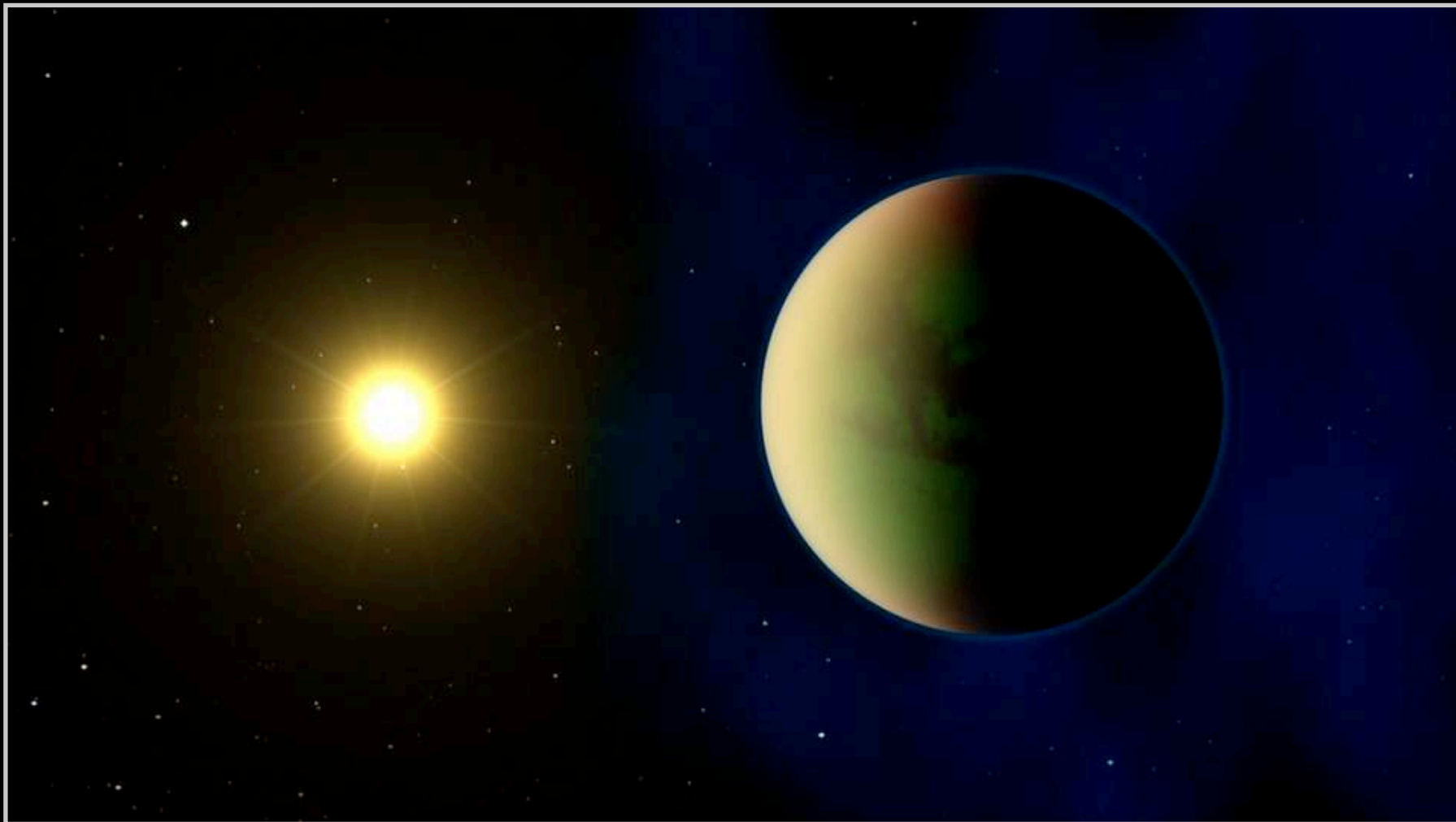
$m \sin(i) \ll \text{mass}$



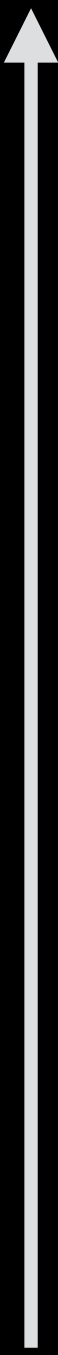
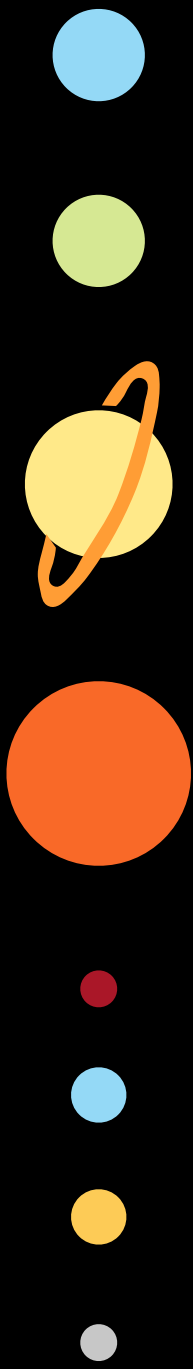
Amount of radiation from star



None of these measurable properties directly relate to surface conditions

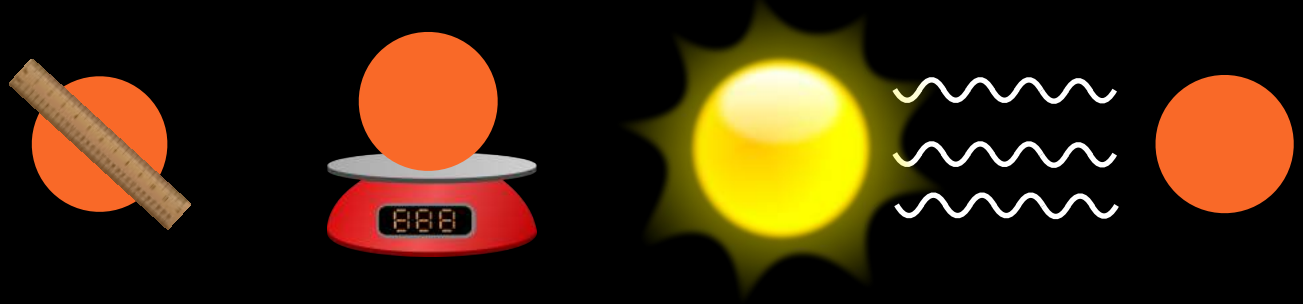


Our next generation of instruments
aim at [atmospheric composition](#)



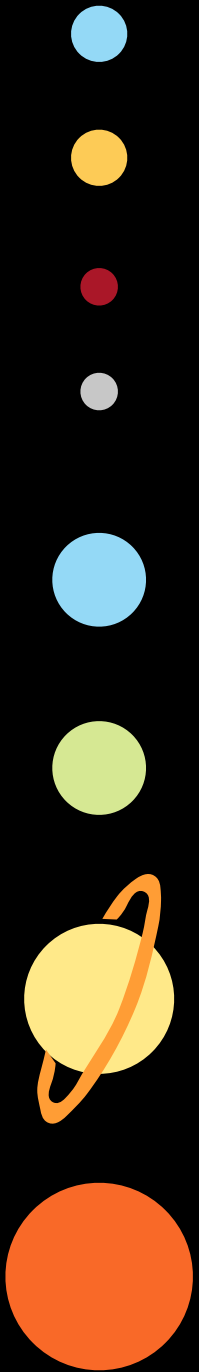
Distance

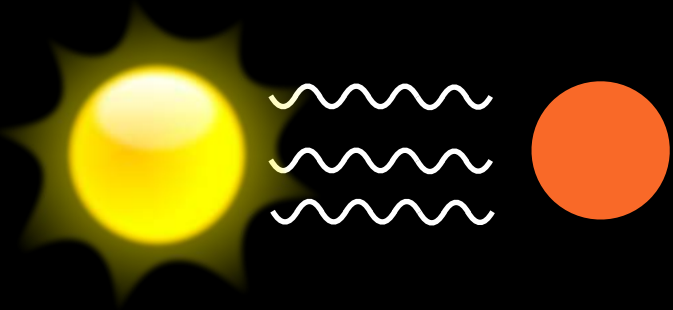
Rank by most interesting target for habitability



... without knowing any surface properties

Priority





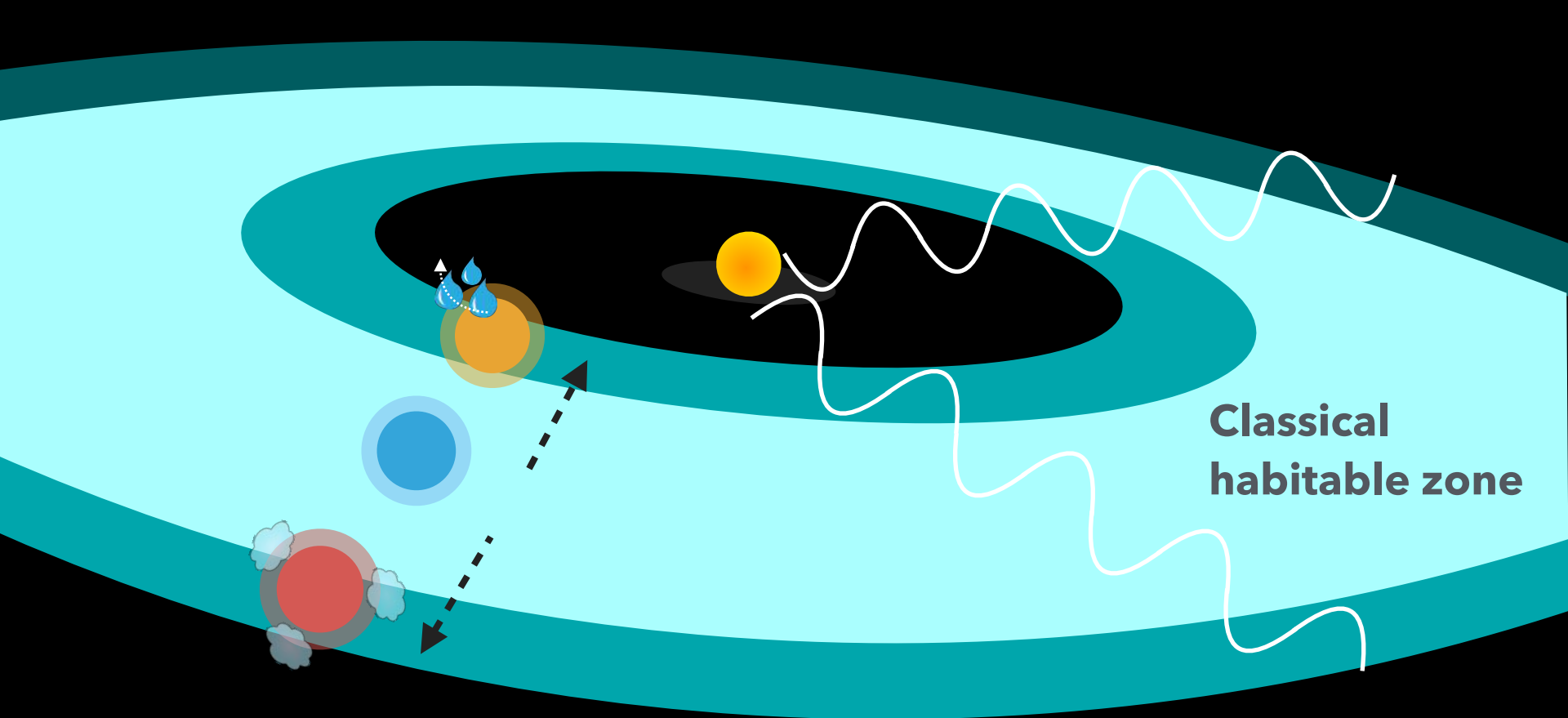
Easiest to recognise Earth-like life
(water & carbon-based chemistry)

Needs to be detectable
(surface water needed)

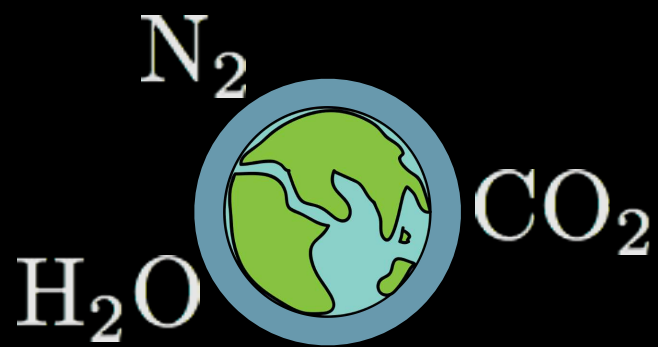
How much insolation does an
Earth-like planet need?

“Classical Habitable Zone”

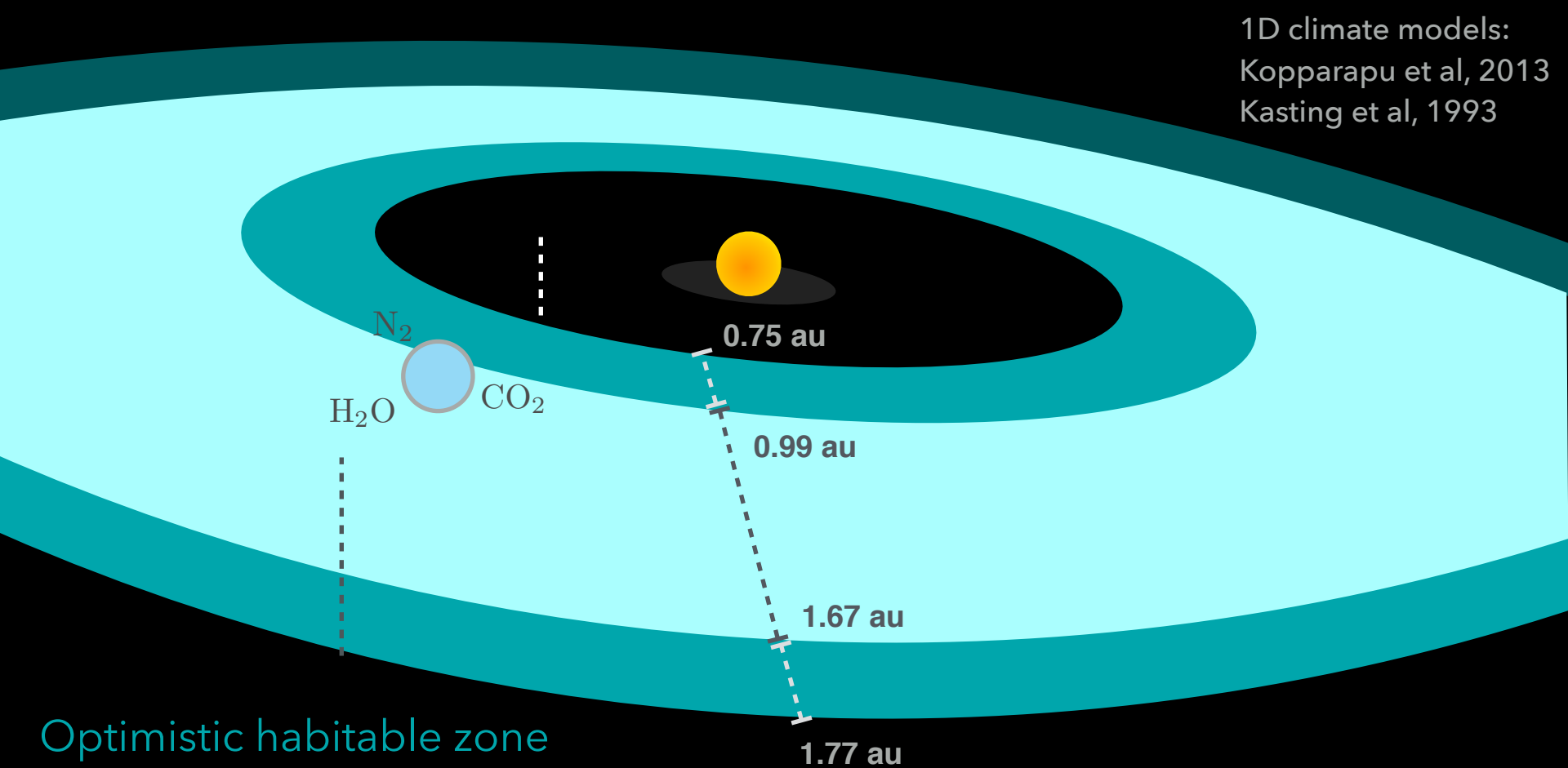




**Classical
habitable zone**

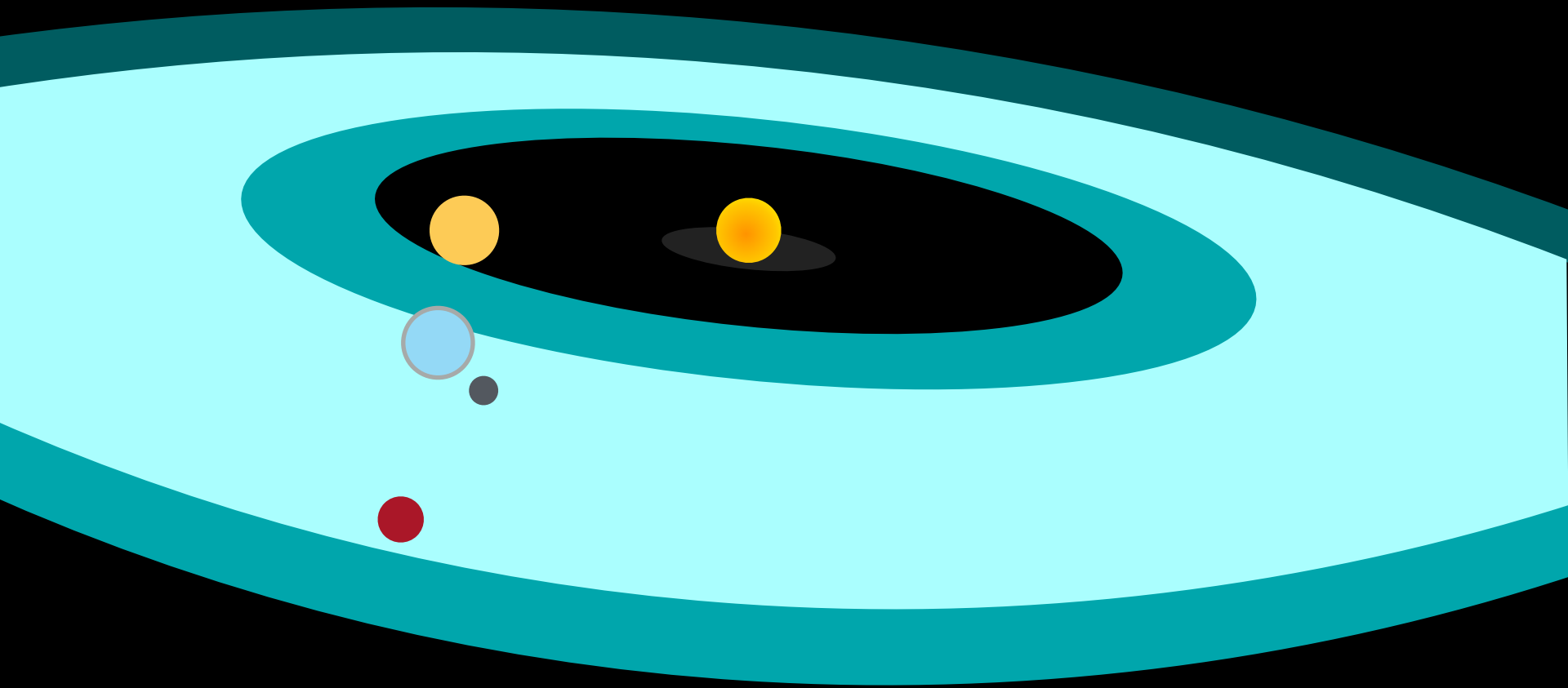


1D climate models:
Kopparapu et al, 2013
Kasting et al, 1993



Optimistic habitable zone

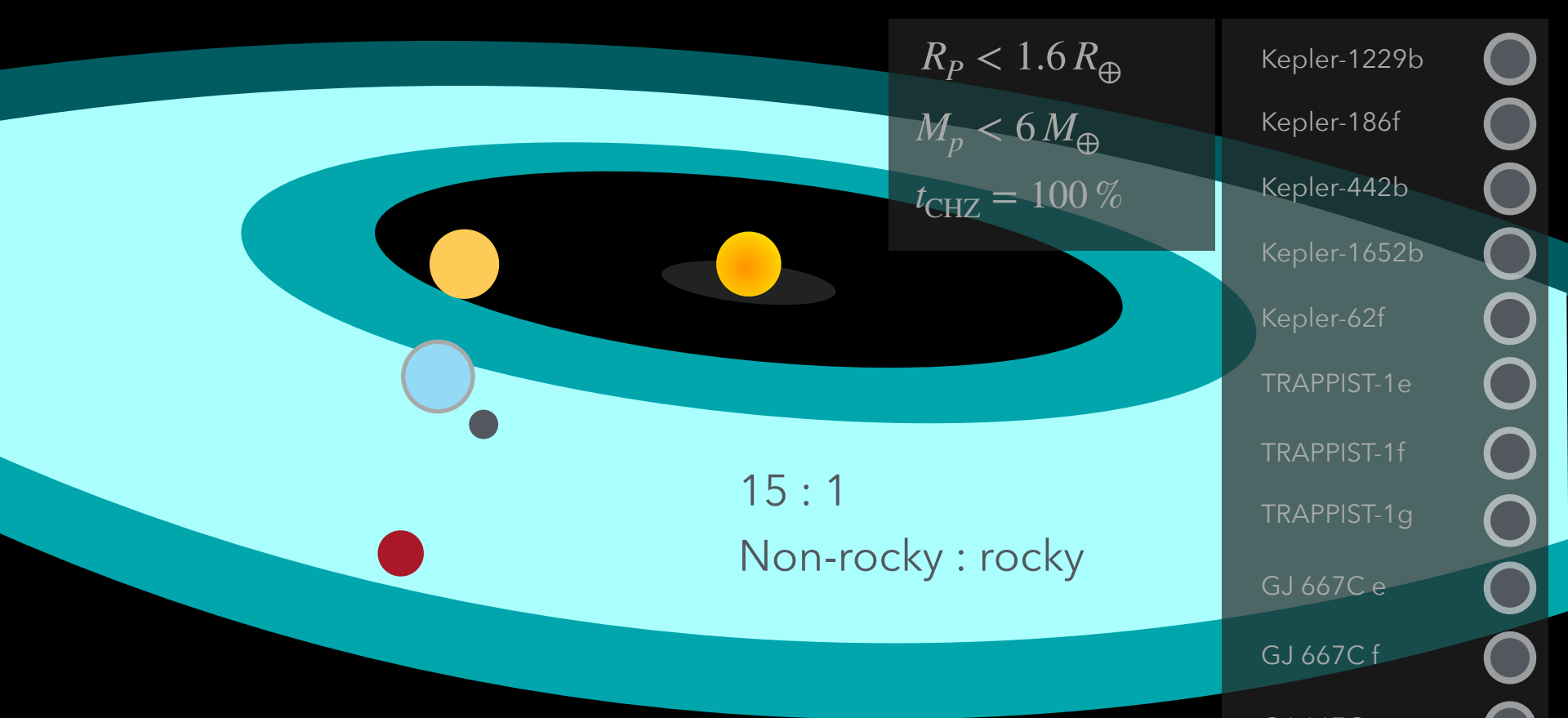
Based on empirical data that
Venus & Mars once had surface
liquid water 1 - 3.8 Gyrs ago



The classical habitable zone is only for an Earth-like planet.

Different planets might have a habitable zone at a different location...

... or not at all.



Are these exoplanets Earth-like?

We don't know.

Can only say:

If we found another habitable Earth-like planet, it would be in the habitable zone.

Conclusions

We've discovered thousands of exoplanets, many of which are similar in size to the Earth.

But at the moment, we have no way of knowing what their surfaces are like (note that the Earth and Venus are both "Earth-sized planets".)

Our next generation of telescopes will be able to detect the atmosphere of these worlds and tell us something about their surfaces for the first time.

The habitable zone is a useful concept for selecting planets for the new telescopes, but it offers no guarantee that a planet is habitable.

Suggested Reading

To see how different even an Earth-like planet might be, try:

earthlike.world



 @earthlikeworld

NASA NExSS 'Many Worlds' blog: www.manyworlds.space

Technical overview of exoplanet
biosignatures:

Yuka Fujii et al,

Astrobiology, Vol. 18, No. 6, 2018

DOI: [10.1089/ast2017.1733](https://doi.org/10.1089/ast2017.1733)

References

“Unique Spectroscopy and Imaging of Mars with JWST”,
Villanueva et al, arXiv:1510:04619

“Habitable zones around main sequence stars”,
Kasting, Whitmire & Reynolds, Icarus, 101, 108-128, 1993

“Habitable zones around main sequence stars: new estimates”,
Kopparapu et al., The Astrophysical Journal, 765:131, 2013

<https://exoplanetarchive.ipac.caltech.edu>